

## Columns for DP-late\_type-H2.xlsx

**Column 1:** Galaxy name

**Column 2:** H<sub>2</sub> mass in M<sub>⊙</sub> within the radius r<sub>25</sub><sup>(1)</sup> derived from CO observations, assuming the constant conversion factor X<sub>CO</sub> = 2 × 10<sup>20</sup> cm<sup>-2</sup> (K km s<sup>-1</sup>)<sup>-1</sup> with ±30% uncertainty from Bolatto et al. (2013). See Casasola et al. (2020) for details<sup>(2)</sup>.

**Column 3:** Uncertainty in H<sub>2</sub> mass M<sub>⊙</sub> calculated as the quadrature sum of the uncertainty on the CO emission line flux and on the X<sub>CO</sub> conversion factor. See Casasola et al. (2020) for details.

**Column 4:** Type of H<sub>2</sub> mass data. Code 0 = detection, Code -1 = upper limit.

<sup>(1)</sup> r<sub>25</sub> is the isophotal radius at which the optical surface brightness falls beneath 25 mag arcsec<sup>-2</sup>.

<sup>(2)</sup> If requested, H<sub>2</sub> masses are also available under the assumption of the metallicity-dependent X<sub>CO</sub> according to the calibration of Amorín et al. (2016). See Casasola et al. (2020) for details.

**Please** do not hesitate to contact us (viviana.casasola@inaf.it) if you intend to use the data for scientific analysis. The following acknowledgement would be appreciated: “This work made use of the H<sub>2</sub> mass data of DustPedia late-type galaxies (Casasola et al. 2020; Davies et al. 2017).”

### References

Amorín, R., Muñoz-Tuñón, C., Aguerri, J. A. L., & Planesas, P. 2016, A&A, 588, A23  
Bolatto, A. D., Wolfire, M., & Leroy, A. K. 2013, ARA&A, 51, 207  
Casasola, V., Bianchi, S., De Vis, P., et al. 2020 A&A, 633, A100  
Davies, J. I., Baes, M., Bianchi, S., et al. 2017, PASP, 129, 044102

### Contact:

Viviana Casasola  
INAF IRA  
Via Piero Gobetti 101  
40127 Bologna, Italy  
viviana.casasola@inaf.it